

This semester: Light (e/m radiation)
 Relativity
 Quantum physics, including atomic and nuclear physics

①

Chapter 33 Electromagnetic (e/m) waves -

OH-
 J Maxwell

Recall from Phys 4B:

\vec{E} = electric field } in empty space
 \vec{B} = magnetic field }

Maxwell's equations - provides a complete description of electricity, magnetism, and light (new this semester)

Gauss's law for E fields
 Chapter 22-2
 of Tipler & Mosca
 Ch. 26 of Wolfson

(1) $\oint \vec{E} \cdot d\vec{A} = \frac{Q}{\epsilon_0}$

(van de Graaf demo)

⇒ properties of the E field

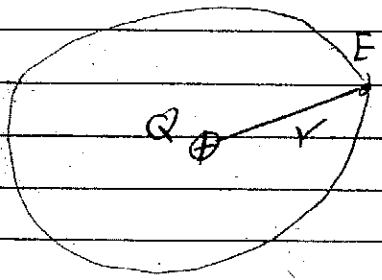
Example - Find E at distance r [E = E(r)]

from a point charge Q:

$F = 4\pi r^2 = \frac{Q}{\epsilon_0}$ since $A_{\text{sphere}} = 4\pi r^2$

$\Rightarrow E = \frac{1}{4\pi\epsilon_0} \frac{Q}{r^2} = k_e \frac{Q}{r^2}$, $k_e = \frac{1}{4\pi\epsilon_0}$

with $F_{\text{Coulomb}} = q\vec{E}$ $\epsilon_0 = 8.854 \times 10^{-12} \frac{C^2}{Nm^2}$
 $\Rightarrow \vec{E}_{\text{Coulomb}} = k_e \frac{qQ}{r^2}$, Coulomb's law.



Most of Phys 4B can be similarly derived from Maxwell's eqs.

Gauss's law for B fields
 Chapter 27-3
 of Tipler & Mosca,
 Ch. 26 of Wolfson

(2) $\oint \vec{B} \cdot d\vec{A} = 0$ (Magnets)

⇒ There are no isolated magnetic charges (monopoles):
 all magnets have both north and south poles.

line integral around perimeter of length s

Faraday's law
 Chapter 28-2 of TAM,
 Ch. 26 of Wolfson

(3) $\oint \vec{E} \cdot d\vec{s} = -\frac{d\Phi_B}{dt}$, where $\Phi_B = \int \vec{B} \cdot d\vec{A}$ where A = area of the loop with perimeter s.

⇒ A time-variable magnetic field (from e.g. a rotating magnet) makes (induces) a current in a loop of wire ⇒ Electric generators
 ⇒ A magnetic field is equivalent to a loop of current

André-Marie
 Ampère's law
 Chapter 27-4
 of TAM,
 Ch. 26 of Wolfson

(4) $\oint \vec{B} \cdot d\vec{s} = \mu_0 I + \mu_0 \epsilon_0 \frac{d\Phi_E}{dt}$, where $\Phi_E = \int \vec{E} \cdot d\vec{A}$ and $\mu_0 = 4\pi \times 10^{-7} Tm/A$

⇒ A time-variable electric field (from a current in a wire, since current $I = \frac{dq}{dt}$) generates a magnetic field ⇒ Electric motors (the opposite of electric generators)

How Maxwell showed light is electromagnetic (e/m) waves - (2)

Observed:

\vec{E} and \vec{B} fields can travel (or propagate) through empty space.

In empty space, containing no sources, $Q = I = 0$ (since $I = \frac{dQ}{dt}$)

In empty space,

Faraday's law: $\oint \vec{E} \cdot d\vec{s} = -\frac{d\Phi_B}{dt} = -\frac{d}{dt} \left(\int \vec{B} \cdot d\vec{A} \right)$

simplified,

this becomes: $E \cdot dx = -\frac{d}{dt} (B \cdot dx - dx \cdot B)$
 $= dA$, since over an area.

$$\Rightarrow \frac{\partial E}{\partial x} = -\frac{\partial B}{\partial t} \quad (33.15)$$

For Ampère's law in empty space, displacement current $\mu_0 I = 0$,

so: $\oint \vec{B} \cdot d\vec{s} = \mu_0 \epsilon_0 \frac{d}{dt} \int \vec{E} \cdot d\vec{A}$

$$\Rightarrow \frac{\partial B}{\partial x} = -\mu_0 \epsilon_0 \frac{\partial E}{\partial t} \quad (33.18)$$

This minus sign should be here: see the rigorous derivation on pp. 879-880 of Serway

Combine eq.s. 33.15 and 33.18 (try it!) and one can get:

$$\frac{\partial^2 E}{\partial t^2} = \mu_0 \epsilon_0 \frac{\partial^2 E}{\partial x^2} \quad (33.19)$$

$$\frac{\partial^2 B}{\partial t^2} = \mu_0 \epsilon_0 \frac{\partial^2 B}{\partial x^2} \quad (33.20)$$

Recall from Chapter 14: These are wave equations.

For any wave traveling along the x-axis:

$$\frac{\partial^2 y}{\partial x^2} = \left(\frac{1}{v^2} \right) \frac{\partial^2 y}{\partial t^2}$$

where y = amplitude of the wave
 and v = speed of the wave, in the +x direction.

For e/m waves, $\frac{1}{v^2} = \mu_0 \epsilon_0 \Rightarrow v = \frac{1}{\sqrt{\mu_0 \epsilon_0}} = 3.00 \times 10^8 \text{ m/s} = c$, the speed of light in empty space

(since $\mu_0 = 4\pi \times 10^{-7} \text{ T}\cdot\text{m/A}$, $\epsilon_0 = 8.854 \times 10^{-12} \text{ C}^2/\text{N}\cdot\text{m}^2$)

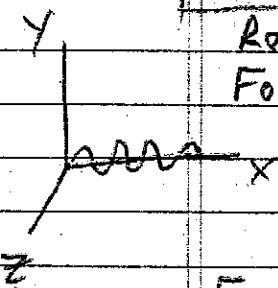
\Rightarrow Light is an electromagnetic wave with speed c , in empty space (in a vacuum)

$$[c = 2.99792 \times 10^8 \text{ m/s}]$$

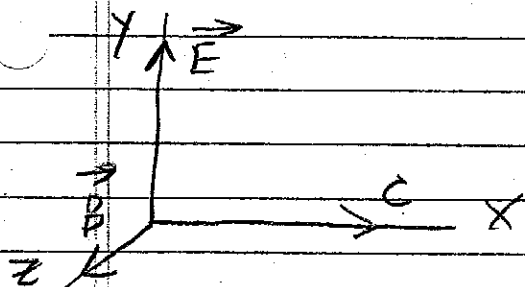
Show off - e/m wave /

do silly dance - demo

with XYZ wire demo



XYZ wire demo



$$\frac{\partial^2 E}{\partial x^2} = \mu_0 \epsilon_0 \frac{\partial^2 E}{\partial t^2} \quad (33.19)$$

$$\frac{\partial^2 B}{\partial x^2} = \mu_0 \epsilon_0 \frac{\partial^2 B}{\partial t^2} \quad (33.20)$$

Review: Chapter 14 (waves)

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These are wave equations. The simplest solutions to them are:

$$E = E_{\max} \cos(kx - \omega t)$$

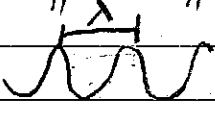
and: $B = B_{\max} \cos(kx - \omega t)$ (Try it!)

demonstrates

These are plane waves,

where E_{\max} = maximum value for the E field

B_{\max} = " " " " " " B " "

λ = wavelength (m) 

k = wave number = $\frac{2\pi}{\lambda}$ (m^{-1} or waves/m)

f = frequency (Hertz or Hz, where $1 \text{ Hz} = 1 \text{ s}^{-1} = 1 \text{ cycle/s}$)

For visible light, $\lambda \sim 4 - 7.5 \times 10^{-7} \text{ m}$ and $f \sim 6 \times 10^{14} \text{ Hz}$.

(Every second, $\sim 6 \times 10^{14}$ light waves go into your eyes.)

From Chapter 14,

to call that: $\lambda f = v$ for any wave.

For EM radiation (light) in empty space, $v = c$,

so: $\lambda f = c$ in empty space.

Angular Frequency $\omega \equiv 2\pi f$.

Notice that: $\frac{\omega}{k} = \frac{2\pi f}{2\pi/\lambda} = \lambda f = c$ ✓

Also, since $E = E_{\max} \cos(kx - \omega t)$

and $B = B_{\max} \cos(kx - \omega t)$,

and $\partial E / \partial x = -\partial B / \partial t$,

$$\frac{\partial}{\partial x} (E_{\max} \cos[kx - \omega t]) = -\frac{\partial}{\partial t} (B_{\max} \cos[kx - \omega t])$$

$$-k E_{\max} \sin(kx - \omega t) = -\omega B_{\max} \sin(kx - \omega t)$$

$$\Rightarrow \frac{E_{\max}}{B_{\max}} = \frac{\omega}{k} = c$$

$$\frac{E}{B} = \frac{E_{\max}}{B_{\max}} = c$$

for EM waves in empty space.

Notice: E and B will be in phase.

Off & headout - Heinrich Hertz & Wilhelm Röntgen

The electromagnetic spectrum

Visible and invisible light are electromagnetic radiation, or the e/m spectrum.

Visible light, which human eyes can see, makes up only a tiny portion of the possible kinds of e/m radiation, the difference between which is their wavelength:

<u>Region of the spectrum</u>	<u>Wavelength (λ)</u>	<u>Typical Sources</u>
Gamma rays	Short!	Nuclear reactions
X-rays	10^{-10} m	Inner-electron transitions
Ultraviolet radiation	0.01 – 0.4 microns (1 micron = 10^{-6} m)	The hottest stars
Visible light	0.4 – 0.78 microns	Outer-electron transitions, and the Sun
Infrared Radiation	0.78-100 microns	Room-temperature objects, including human bodies and planets
Microwaves	millimeters	
Radio waves	centimeters & longer	The very cold gas between the stars

The properties of electromagnetic (e/m) radiation

In 1887, Heinrich Hertz experimentally confirmed Maxwell's electromagnetic theory, by being first to transmit and receive radio waves.

In 1895, Wilhelm Röntgen serendipitously (accidentally) discovered X-rays.

They were named X-rays because their nature was unknown, at first. (The symbol X denotes an unknown.)

In a series of experiments, Hertz showed that radio waves are e/m radiation, with wavelengths over 1000 times *longer* than visible light.

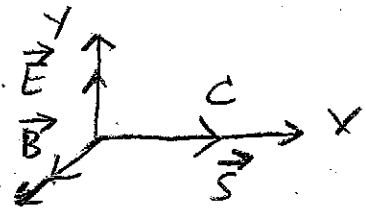
In a similar series of experiments, Röntgen showed that X-rays are also e/m radiation, with wavelengths 1000 times *shorter* than visible light.

Both series of experiments did this by showing that radio waves and X-rays have these properties:

- They have speed $v = c$, the speed of light, in empty space.
- They are *transverse* waves, with $\vec{E} \perp \vec{B} \perp$ the direction of travel.
- They have $E/B = E_{\max}/B_{\max} = c$ in empty space.
- They *reflect* (Chapters 34 and 35).
- They *refract* (Chapters 34 and 35).
- They *interfere* with other waves (Chapter 36).
- They *diffract* (which means bend around corners) (Chapter 37).
- They can be *polarized* (which means they can have E fields only in a specific direction) (Chapter 37).

Visible light does all this too!

So do all other forms of e/m radiation, including infrared and ultraviolet radiation, microwaves, etc.



John Henry Poynting
1852-1914
Student of Maxwell's

E/M waves carry energy -

The energy flux (or flow) is given by the Poynting vector:

$$\vec{S} = \frac{1}{\mu_0} \vec{E} \times \vec{B} \quad \text{Units: } W/m^2 \text{ (Power/Area)}$$

Example - At Earth, we get an average $|\vec{S}|_{avg} = S_{avg} = I$ (intensity)
 $= 1340 W/m^2$ of sunlight, from the Sun. (The text says "x1,000 $\frac{W}{m^2}$ ")

For plane waves, $|\vec{E} \times \vec{B}| = EB$,

$$\text{so } |\vec{S}| = \frac{EB}{\mu_0} = \frac{E_{max}^2}{\mu_0 c} = \frac{B_{max}^2 c}{\mu_0} \quad \text{since } \frac{E}{B} = \frac{E_{max}}{B_{max}} = c$$

Also, Intensity (W/m^2) = $I = \frac{E_{avg}^2}{\mu_0 c} = \frac{E_{max}^2}{2\mu_0 c}$ since $\langle \cos^2(kx - \omega t) \rangle = \frac{1}{2}$

I falls off with r^2 ,

since Power = $I_1 A_1 = I_2 A_2$
 $= I_1 4\pi r_1^2 = I_2 4\pi r_2^2$

so $\frac{I_1}{I_2} = \frac{4\pi r_2^2}{4\pi r_1^2}$ and $\boxed{\frac{I_1}{I_2} = \left(\frac{r_2}{r_1}\right)^2}$

so $\frac{1}{2\pi} \int_0^{2\pi} \cos^2(kx - \omega t) dt = \frac{1}{2}$

Example - Cassini spacecraft at Saturn
 $r = 10 \text{ AU}$
 $r_{Earth} = 1 \text{ AU}$
 $\frac{I_{Saturn}}{I_{Earth}} = \left(\frac{r_{Earth}}{r_{Saturn}}\right)^2 = \left(\frac{1}{10}\right)^2$
 $I_{Saturn} = \frac{1}{100} I_{Earth} = \frac{1340 W/m^2}{100}$
 $= 13.4 W/m^2 \checkmark$

Energy density - Joules/ m^3

Recall from Phys 4B:

Electric energy density $U_E = \frac{1}{2} \epsilon_0 E^2$ (33-28) (25-15)

Magnetic energy density $U_B = \frac{1}{2\mu_0} B^2$ (33-29) (31-14)

For an e/m wave,

$$U_E = U_B = \frac{1}{2} \epsilon_0 E^2 = \frac{B^2}{2\mu_0} \quad \text{since } \frac{E}{B} = c$$

Total instantaneous energy density

$$U = U_E + U_B = \epsilon_0 E^2 = B^2 / \mu_0$$

and $\frac{1}{c^2} = \mu_0 \epsilon_0$

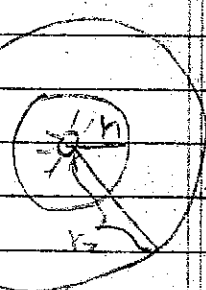
Total average energy density

$$(U_{total})_{avg} = \epsilon_0 (E^2)_{avg} = \frac{1}{2} \epsilon_0 E_{max}^2 = \frac{B_{max}^2}{2\mu_0} \quad \text{since } \langle \cos^2(kx - \omega t) \rangle = \frac{1}{2}$$

$I = |\vec{S}|_{avg} = (U_{total})c \Rightarrow$ Energy density travels with speed c ,
 in empty space.

Units for energy density:

$$U_{avg} = S_{avg} / c = \frac{(W/m^2)}{m/s} = \frac{W \cdot s}{m^3} = \frac{J}{m^3} \quad \text{(since } 1 \text{ watt} = 1 \frac{\text{Joule}}{s} \text{)}$$



- $\rho(\text{air}) = 1.225 \times 10^{-3} \frac{g}{cm^3}$
- $n(\text{air}) = 2.5 \times 10^{17} \frac{\text{molecules}}{cm^3}$
- $\rho(\text{H}_2\text{O}) = 1 \frac{g}{cm^3}$
- $\rho(\text{Fe}) = 7.9 \frac{g}{cm^3}$
- $\rho(\text{Pb}) = 11.4 \frac{g}{cm^3}$

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E/M waves carry momentum -

Radiation pressure $P = |\vec{S}|/c$ for 100% absorption
(black surface)

$P = 2|\vec{S}|/c$ for 100% reflection
(white surface)

$$\text{Since } |\vec{S}| = \frac{1}{\mu_0} |\vec{E} \times \vec{B}| = \frac{w}{m^2}$$

$$P = \left(\frac{W}{m^2}\right) / \left(\frac{m}{s}\right) = \frac{\text{Newtons}}{\text{Area}} = \text{Pressure } \checkmark$$

OH - Comet tails

OH - Solar sail